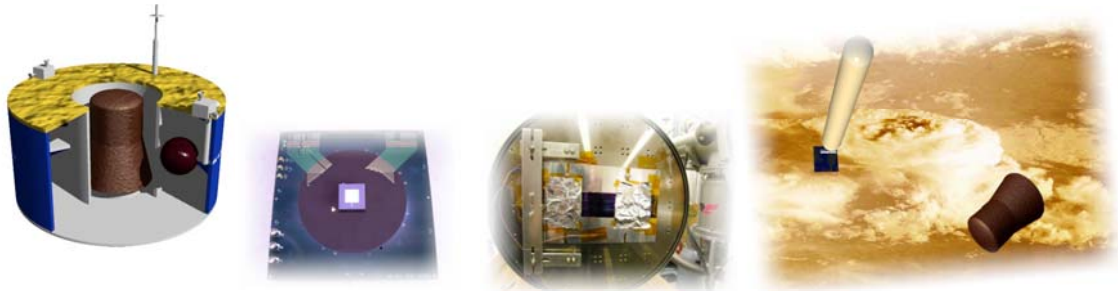


IPPW-6

Research Activities on Venus Atmosphere Balloon Observation Mission



By Tetsuya YAMADA ¹⁾, Kazuyuki HIROSE ¹⁾, Koji TANAKA ¹⁾, Hiroshi TAKEUCHI ¹⁾,
Naoki IZUTSU ¹⁾, Kazuhisa FUJITA ²⁾, Nobuaki ISHII ¹⁾

Japan Aerospace Exploration Agency

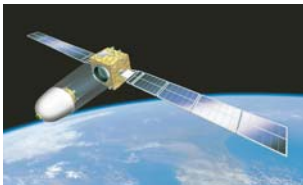
1) Institute of Space and Astronautical Science

2) Institute of Aerospace Technology

1

Be a Dawn of Japanese Planetary Entry !

Reentry Researchers in ISAS
have been engaged in...



USERS/REV

dedicated to μ G Experiment
ISAS cooperate with USEF on Research
Activities

Launched Sept/2002

Recovered May/2003



EXPRESS

Launched Jan/1995

Reentry Tech. Acquisition

What's Next ?

we would answer

Planetary Entry ! Begin with Venus !!

HAYABUSA

Asteroid Sample Return

Launched April/2003

Arrive at Asteroid Aug/2005

Return June/2007



DASH

Launched Mar/2002

-Hyperbolic Velocity

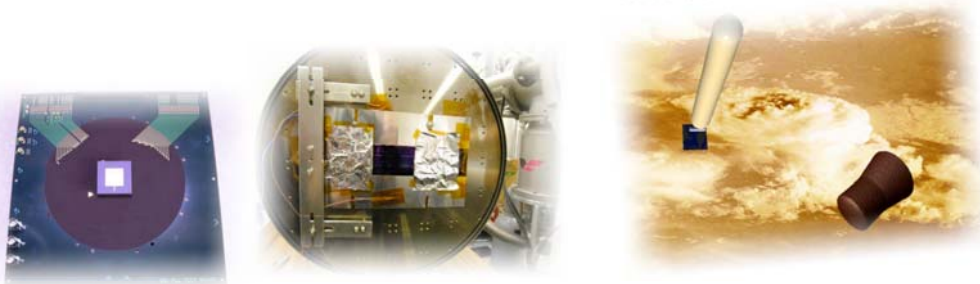
-Precursor for M-C

2

Contents of the Presentation

Research Activities on Venus Atmosphere Balloon Observation Mission

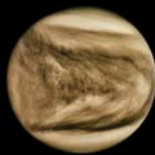
- Brief Introduction of Target Mission Concept
- Recent Status including External Relations
- Brief Outline of the Subsystems
with Recent R&D Activities for Critical Issues
- Future Work and Schedule addressed



Planet Venus and Probe Explorations

Venus : Twins of Earth but Mysterious Planet in the Solar System

- 0.72 AU from the Sun
- R=6052 km (95% earth), 0.815 Earth Mass
- Rot. Period = 243 days (reverse)
- T=460°C, P=90 atm (Surface)
- T=200°C (@H=35km)
- Strong Equatorial Wind
- Sulfuric acid Cloud (H47-70km)



※Observations under clouds is
Significant from Scientific Point of View

Venus

Great Historical Probe Missions

Venera 4-16
(USA, 1967-84)

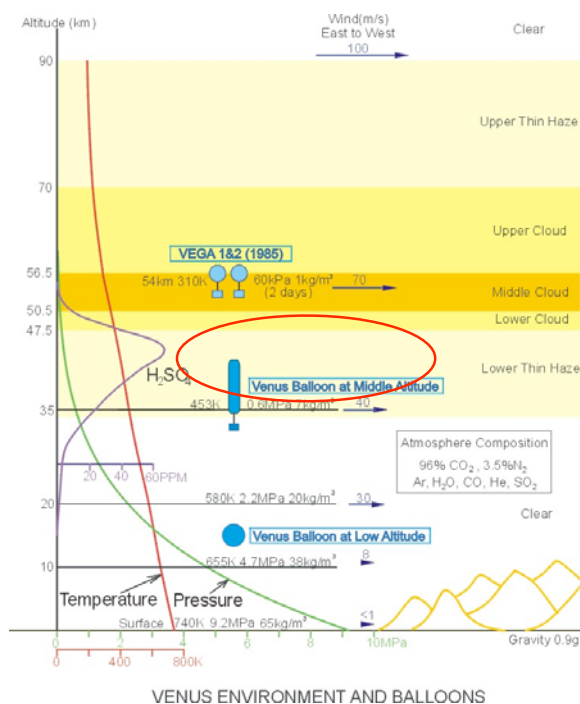
Mariner 2-10 (USA,
1962-78)

Pioneer-Venus,
(USA, 1978)

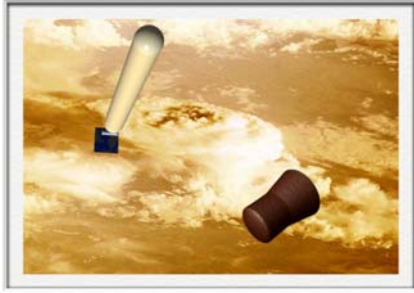


historically, lots of probes
in USA, USSR

VEGA
(USSR/FR 1984)



Long-term Observation under the Clouds



Long-term Observation of Low-Altitude Venus Atmosphere by Tracking of water-vapor Balloon

- Target Altitude : H=35km (under the Clouds)
- Mission Period : beyond 2 weeks' observation

1) Scientific Significance

Long Term Observation under the Cloud(H70-47km) will reveal....

- Mechanism of the Strong Equatorial Wind, N-S circulation
(Internal Gravity Wave, Turbulence, Structure of Vertical Wind)
- Concentration of Aerosol (unknown particle) : optional due to the small weight budget
- Precise Mapping of the Venus Surface($\lambda \sim 1\mu\text{m}$) : optional

2) Engineering Significance leading to future Planetary Exploration

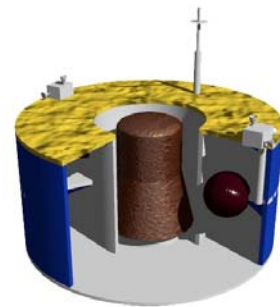
The mission is meant to be

- **Dawn** in Aerothermodynamic Technology on Atmospheric Entry
for Future Outer Planets' Exploration in JAPAN.
- **Demonstration** of the High Temperature Electronics in the Hot Venusian Atmosphere.
- Planetary Long Term Observation by Balloon itself is of significance

5

Recent Status including External Relations

Long-term Observation of Low-Altitude Venus Atmosphere by Tracking of water-vapor Balloon



1) Proposal in ISAS / JAXA

Authorized as a **pre-phase-A WG** for "Future Small Scientific Sat"

- Not Selected to Proceed to phase-A for 1st launch of Small Solid Rocket in 2012FY

2) International Collaboration

Applied to COSMIC Vision 2018 of European Science Academy

- "JAXA low-altitude balloon" : one option of EVE(European Venus Explorer)
- **Passed First Selection but** not selected as Final 8 Candidates.

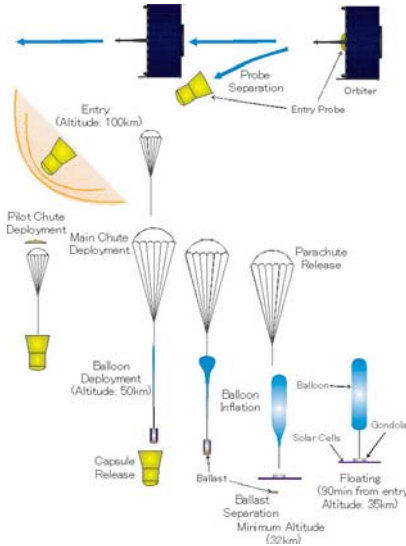
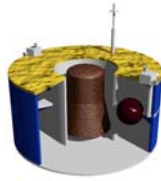
**Never Give up and
Apply again !!
Improving Technical Readiness Levels.**

6

Technical Issues associated with the Mission

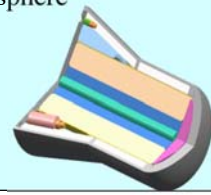
Weight : 160 kg
Size : $\phi 1.4 \times H 0.7$ m
Cap (BALN,PI) : 35kg

Direct Entry
 $V_0 = 11.5$ km/s, $\gamma \sim 15$ deg



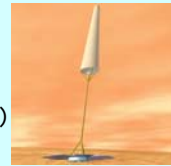
1) Venus Entry Capsules

- High Speed Entry to CO₂ Atmosphere
 - Flight Environment Prediction
 - Thermal Protection Design
- TPS Development and Facility
 - High Enthalpy CO₂ Generation
- Descent System



2) Water-vapor Balloon

- Long-term Observation (2 weeks)
- Multi-Layered Balloon Film (Gas-Barrier, Lightweight, High Strength)
- Efficient Heat-exchange and Inflation



3) Tracking of the Probe

- narrow-band DDOR under High Temperature

4) High-Temperature Electronics

- High Temperature Electronics over 180°C (Solar Battery Cells, Oscillator/Transmitter, Other SOI devices)



7

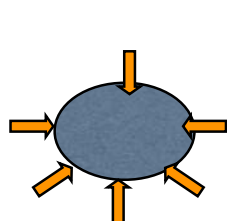
Water-vapor Pressure Balloon

Water-Vapor Balloon System makes 2 kg Bus+PI fly at H=35km .

**For Successful Inflation,
Efficient Heat-convection from the
atmosphere is Important !**

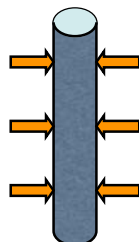
**Water-Mass / Surface Area Ratio is a Key
Parameter**

Low M/S is desirable



Pumpkin-type

- Hi Buoyancy
- light Weight



Cylindrical-type

- large heated surface
- needs large film
- => weight penalty

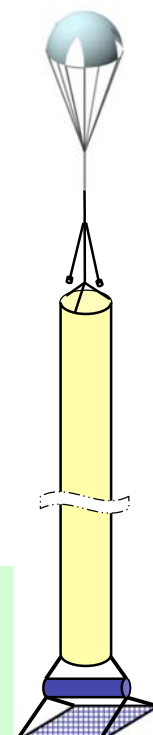
**18 m Long
Balloon**

$\Phi 0.35\text{m} \times 18\text{m}$
@ Full-expansion

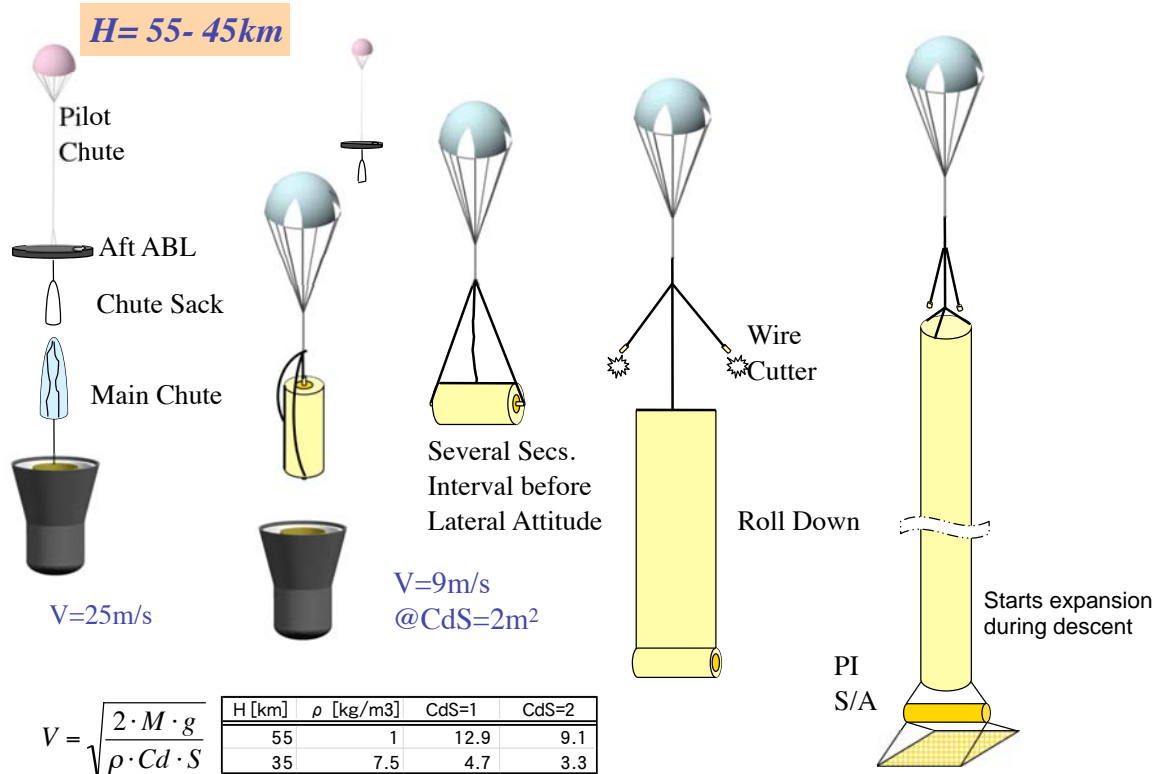
Weight Allocation

Bus+ PI = 2 kg
Balloon = 3.2 kg
Water = 4.8 kg

Slow Descent
by Parachute
until Full-Inflation



Balloon Deployment Sequence



9

Balloon Release Altitude from 55 to 45 km

Constraints for Release Altitude

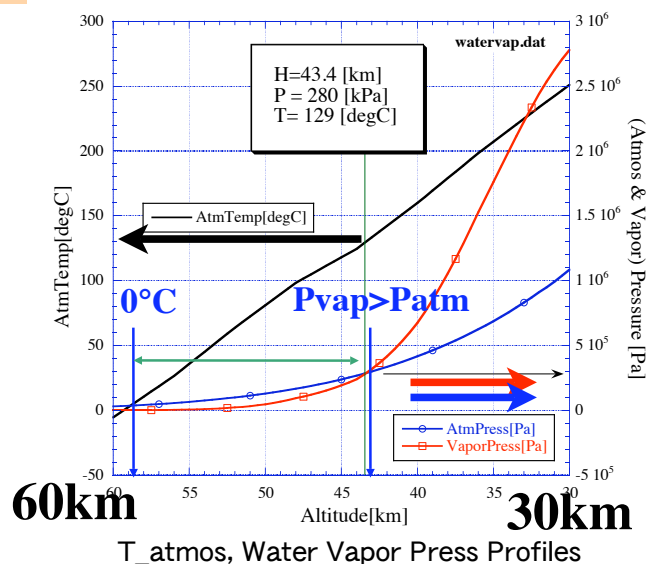
- Liquid Water must not be frozen
- avoid Inflation inside Capsule

• BALN is Never Released
@H > 55 km (T ~ 0°C)
so as not to Freeze the BALN Water

• BALN can Inflate (@H=43 km)
Vapor Press > Atmos Press
→ BALN released @H=45km

Venus Atmosphere				
Z [km]	T [degC]	ρ [kg/m³]	P [Pa]	As [m/s]
59.3	0	0.62	4289	260.2
47.7	100	2.4	4091	299.6
37.9	180	5.97	4073	327.8
35.0	207	7.5	4074	336.7

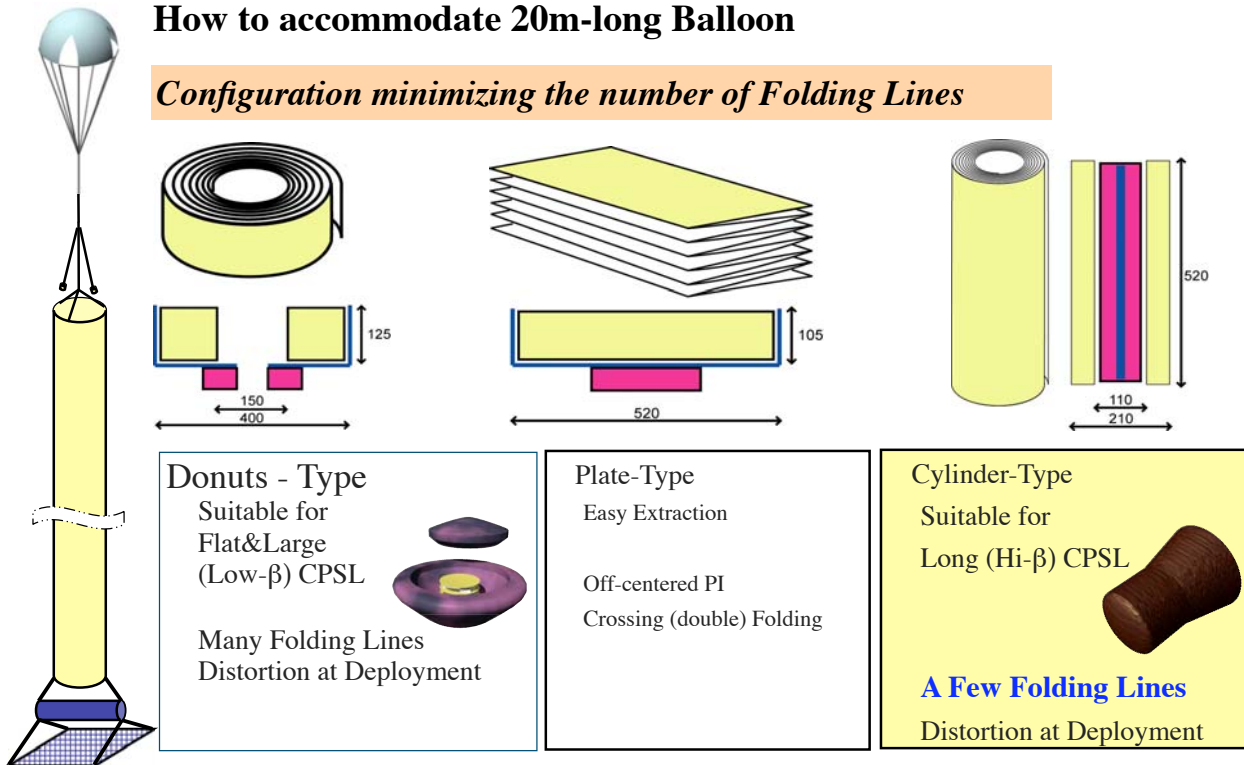
Balloon Release Altitude 55km ~ 45km



Trade-off Studies on Balloon Accommodation

How to accommodate 20m-long Balloon

Configuration minimizing the number of Folding Lines



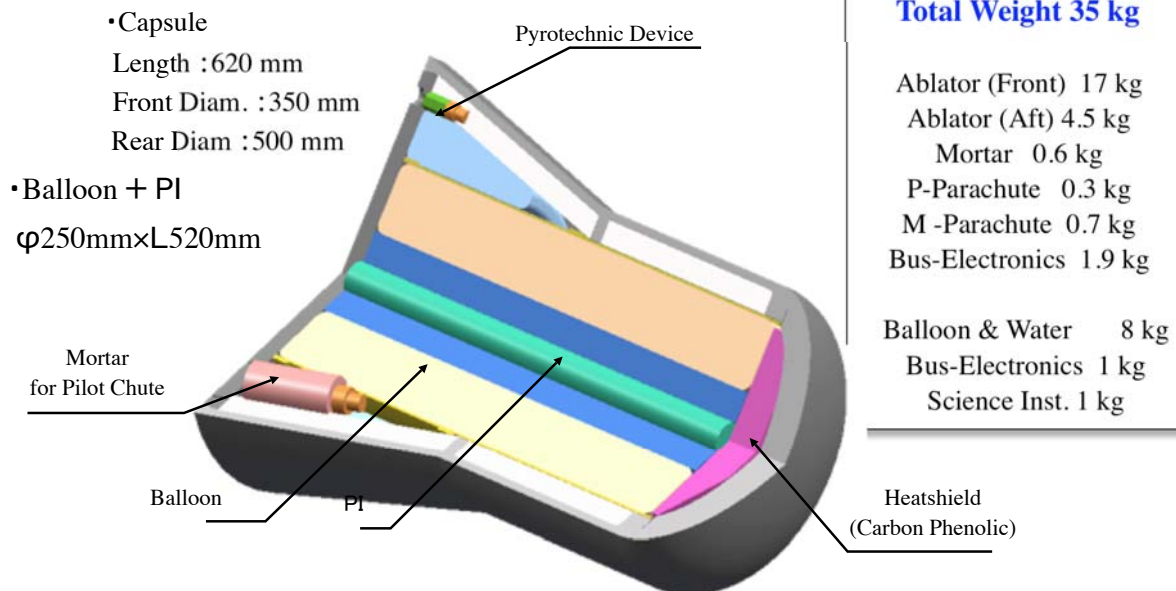
11

Venus Entry Probe (Hi-Ballistic Coeff-type)

For protecting the inside against the soak-back heat from the heatshield

Fast Descending Flight is required for the Capsule.

Balloon accommodation => Hi-Ballistic Coeff. Capsule



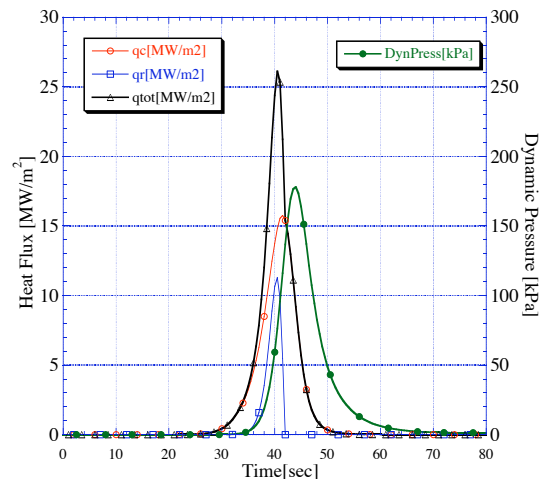
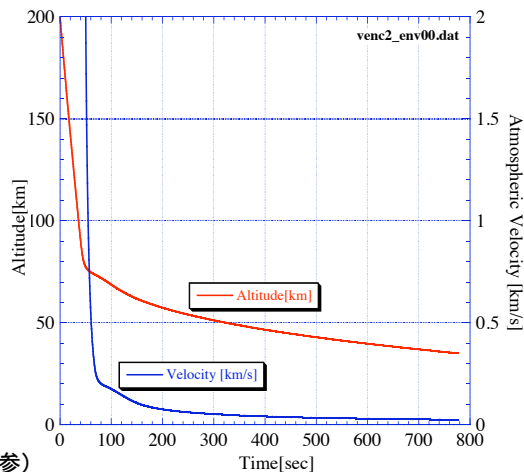
Flight Environment of the Capsule

Max. Heat Flux expected

Z : 89.4 km
V : 10.62 km/s

max. $q_r = 12 \text{ MW/m}^2$
max. $q_c = 16 \text{ MW/m}^2$
max. dyn. Press = 180kPa

Atmos. Pressure : 45.2 Pa
Atmos. Temperature : 173.8 K



参)
dyn.P @ 100km/h
= 500Pa = 5/1000 気圧

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Venus Entry and TPS Development

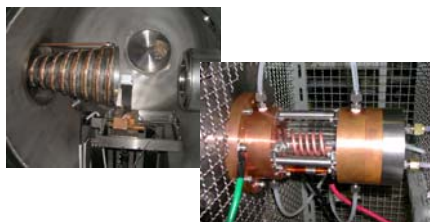
TPS Development Scenario

Ablator Material Requirements
• Protect Capsule
• against High Heatflux
• Lightweight



Induction-coupled Plasma Generator (ICPG) 10 kW

※ High enthalpy CO₂ heated by 13.56MHz RF
• Simulate Thermochemical Aspect
of Reaction between CO₂ and Ablator



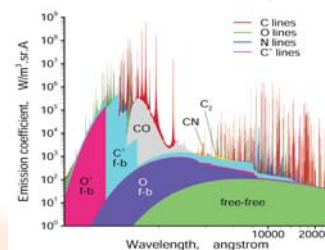
Arc windtunnel : 1 MW

=> can generate Hi-Enthalpy Air but not CO₂
due to Carbon deposit to the Electrodes
but very useful in
• Material Thermal Strength Test in High Heat Flux upto 15MW/m²
• Temperature Profile Data

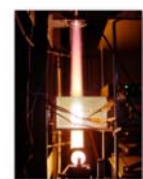
Thermochemical Basic Data

- Reaction Rate Measurement in Hi-Enthalpy CO₂
- Numerical Simulation
- CFD Analysis

utilizes the Ground Simulation Result to predict Flight Environment and Thermal behavior of the Ablator



Exmpl : Radiative Heat Flux Analysis

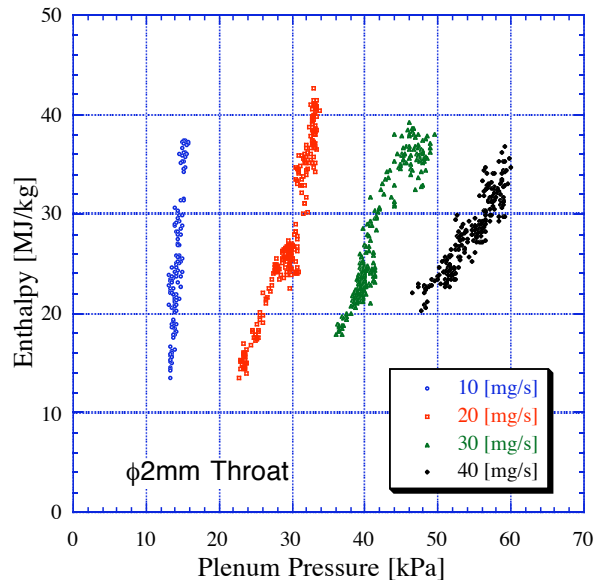
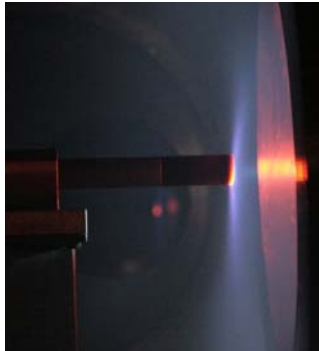


触媒性計測装置

Venus Entry and TPS Development (1/2)

Characterization of ICPG and Material Heating Test are now carried out ...

- 40 MJ/kg Enthalpy Accomplished
=> useful for Thermochemical Data of Marial/CO₂ Reactions
- Due to the nozzle installed on ICPG
Plenum Pressure ranges upto 50 kPa and
Impact Pressure ranges upto 5 kPa

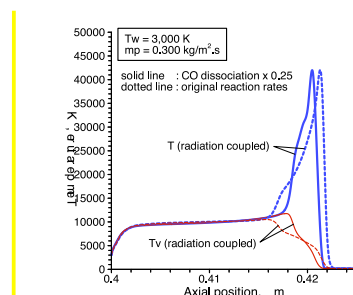


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Thermochemical Aspect of Venus Entry

Improvement in Flight Environments Assessment

CO thermal relaxation & dissociation has great effect on both q_c & q_r



Modification	$\dot{m}_p = 0.3 \text{ kg/m}^2\text{-s}^a$			
	Uncoupled		Coupled	
	q_c	q_r	q_c	q_r
CO dissociation rate $\times 0.25$	4.6	17.5	6.0	12.0
rate of reaction 23) from Ref. 12	4.6	11.3	5.3	8.2
rate of reactions 22) and 23) from Ref. 12	4.6	11.3	5.3	8.2

^a $T_W = 3,000 \text{ K}$. ^b $T_W = 300 \text{ K}$.

Nominal 4.6 11.2 5.4 8.0

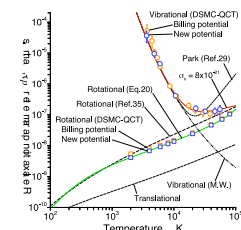
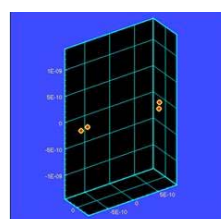
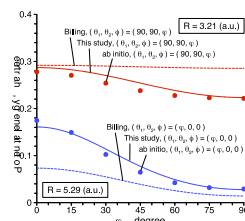
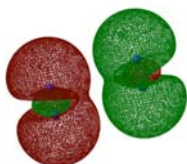
Development of improved models for CO relaxation & dissociation

MO analysis

High accuracy PES

QCT collision analysis

CFD model



16

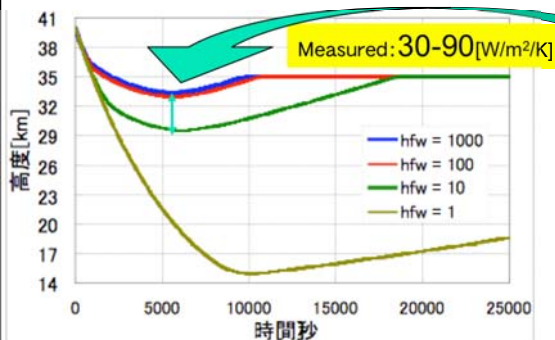
Inflation Analysis of Balloon Film

Analysis of Heat-exchange Process

Heat Transfer Rate between Film – “Water sheet” is important

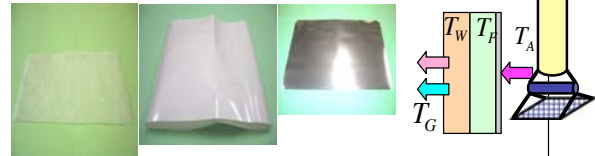
High Heat Transfer Rate
 → Fully-expanded at higher altitude
 → Burst out due to insufficient film strength

Low Heat Transfer Rate
 → Late Inflation, and Overshoot to low Z.
 → exposed to unexpected Hi-Temp.

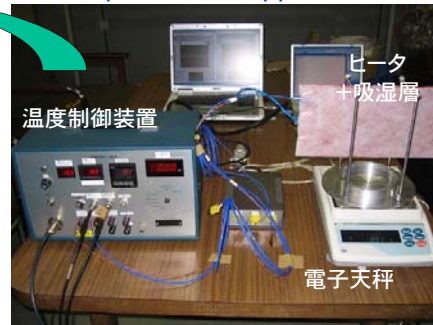


Heat-Convection Measurements

Measurement of Gas Permeability and Heat Convection



experimental apparatus



17

Balloon Expansion Simulation

【Lessons learned until 2006FY】

- Fabrication of Subscale Model (D=0.159m, L=1.88m)
- Expansion Simulation in Hot-Airflow (140°C) Cavity

=> Successful Expansion in 180 sec.
 within dispersion of prediction
 → inflexibility problem of the film
 was revealed during the experiment !



【Research Activities in 2007FY】

Balloon Film made of Liquid-crystal Polymer (LCP)

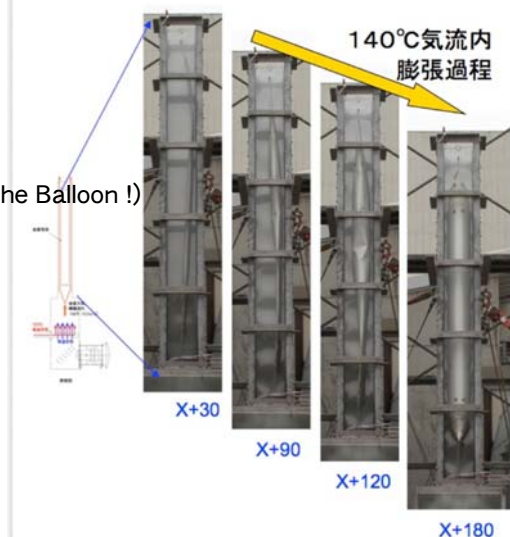
- Good Performance in Gas-Barrier Characteristics
- Manufactured in Cylindrical-shape Inflation (desirable for the Balloon !)
- Drawback : Film has hard/ poor-flexibility, hard to be accommodated

Change of Resin

make the film more flexible, easy to treat.

: from PolyPlasitc → Sumitomo Chemicals

- Strength in High-Temperature (now 103% of Pa)
- Flexibility
- Corrosion due to High-Temperature Water



18

Tracking of Balloon by DDOR

DDOR (delta Differential One-way Ranging)
normal VLBI(Cont. Wave like QUASAR)

$$\text{Sensitivity} \propto (\text{Band Width})^{0.5}$$

Wide Band detection is identical to large
Transmittance Power



Narrow-band VLBI

(for spacecraft and probes)

$$\text{Sensitivity} \propto (\text{Band Width})^{-0.5}$$

because of transmitter power limitation

Integration Time $\sim 100\text{sec}$

(Wind Speed @H35km =30m/s

=> Spatial Resolution \sim several km

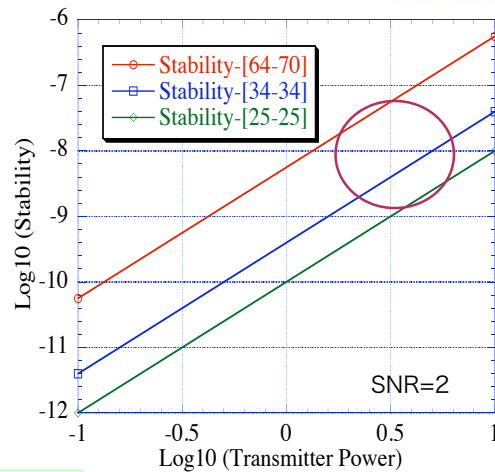
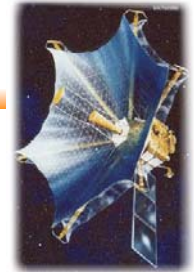
Scientific Positioning Request ($\sim 2,30\text{km}$)

depends on Transmitter Frequency Stability

Transmit Power \sim Several W $\rightarrow 10^{-9}$ Stability

VEGAは1.7GHz,6.5MHz離れた2波長,送信パワ 5W,発生電力20W

VLBI・Radio-Telescope
HALCA (M-B) / Astro-G



金星-地球間距離 (最大:1.7AU)

1AU=1.495e11 [m]

参)月・地球 38万km = 3.8e8 [m]

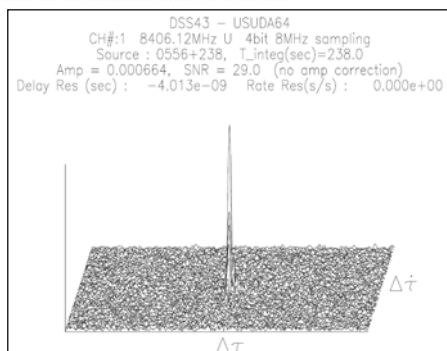
金星上距離 100km = 7e-7 rad

DDOR Experiment by using Hayabusa S/C

- JPL/NASA, ESA and JAXA collaborative Experiment.
- Δ DOR Signal sent from Goldstone (JPL) (± 1 MHz)、Return at Hayabusa
- Received at Goldston(JPL),Canberra (JPL)、Usuda, Kashima.

Canberra 70m —Usuda 64m

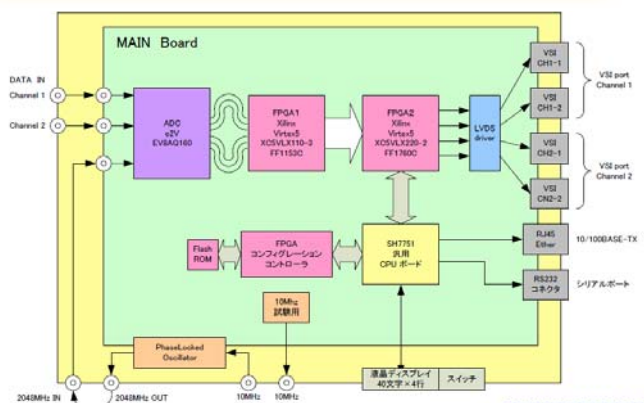
Detected Quaser Flinge



- Fluctuation of venus atmosphere can be cancelled by using Quaser in the vicinity of Venus.



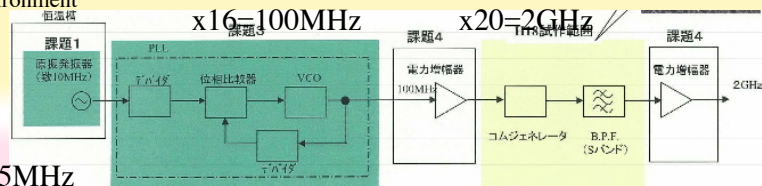
VLBI Data Acquisition System (ADS-3000)
for narrow band detection



High-Temperature Electronics

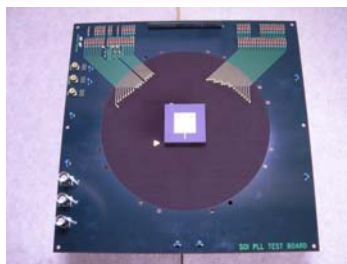
High-Temperature Electronics are Key Technology for Tracking Low-altitude Venus Balloon.

- Quartz Oscillator : Stability beyond 10^{-9} by means of Appropriate Crystal-Cutting and Temperature Control.
- Solar Cells operable in 200 degC environment
(Though 200 degC in operation)

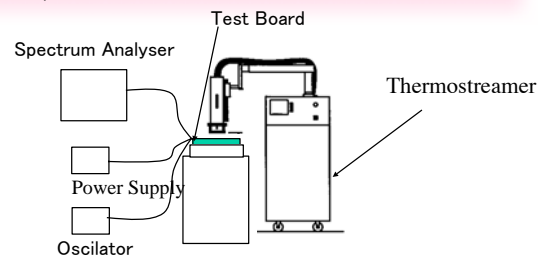


【Demonstrated and Characterized】

- Frequency UpConverter (100MHz to 2 GHz) in 190 degC (2006)
- PLL upconverter (2007)
- Amplifier and Oscillator (planned in 2008)



Test Board (IC is centered)



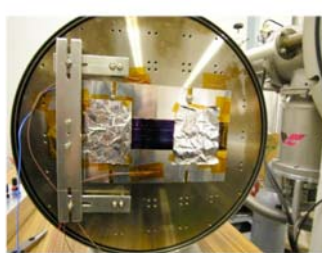
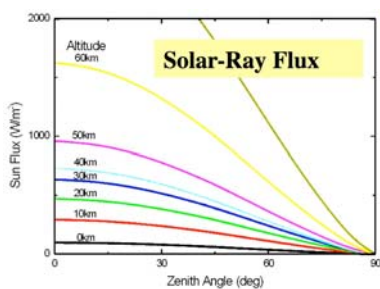
Experimental Setup

21

Thin-Film Solar Cells in High-Temp (1)

Temperature Characteristics of Thin-film Solar Battery Cell were obtained;

- Tandem-type Amorphous-Silicon Solar Battery Cells formed on Polyimide Film.



Experimental Configuration

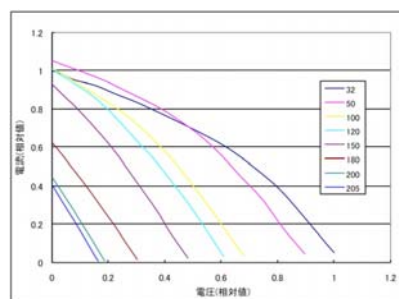


Experimental Apparatus

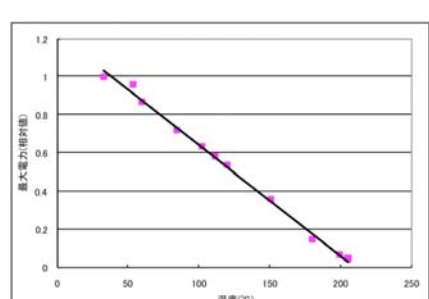
- Temperature : 180°C
- Input Irradiance : 200W/m²

Test Cells:
by Fuji Electric Systems co.
(富士電機システムズ)

Results:
Temperature Characteristics:
Voc: -0.46%/°C
Pmax: -0.56%/°C



Temperature Profiles of V-I Characteristics



Time Profile : Max Output Power

22

Thin-Film Solar Cells in High-Temp (2)

Spec of a TEST Module (12Cells Series)

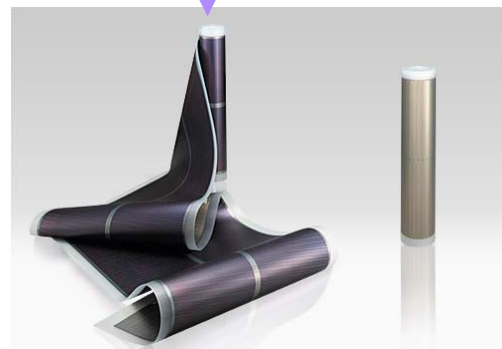
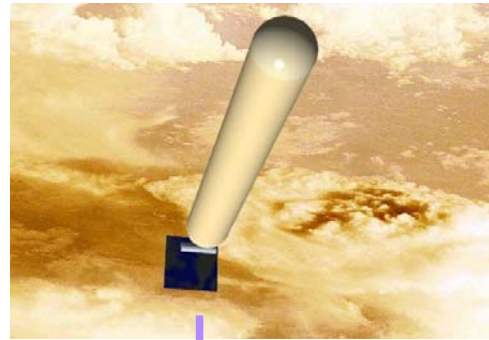
- Max Output Power : 2W
- Max Voltage : 13V
- Dimension : 170×240mm
- **Conversion Eff. : about 0.7% (@180degC)**
- 1μm Solar Cells on 50μm Polyimide Film

Design Performance on the Venus estimated by the above results

- Operable Temperature : 180°C
- Input Irradiance : 200W/m²
- ⇒ **Generated Power : about 1.5 W/m²**

Research Issues

- forming Cells on Balloon film ?
- Surface Protection Film with Anti-Acid Characteristics
- Adhesive bonding



a-Si薄膜太陽電池:富士電位システムズ[FWAVE]

23

Research and Development Schedule

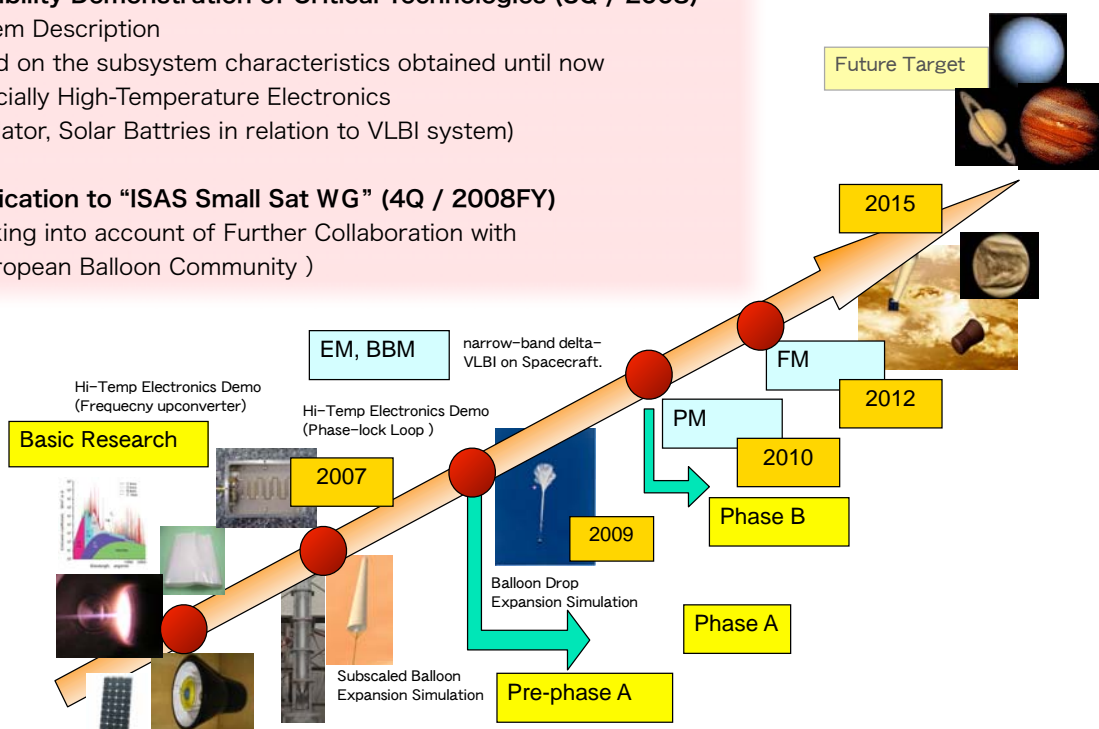
Schedule

• Feasibility Demonstration of Critical Technologies (3Q / 2008)

System Description
based on the subsystem characteristics obtained until now
especially High-Temperature Electronics
(Oscillator, Solar Batteries in relation to VLBI system)

• Application to "ISAS Small Sat WG" (4Q / 2008FY)

(Taking into account of Further Collaboration with
European Balloon Community)



24

Summary

Research Activities on Venus Atmosphere Balloon Observation Mission

- Introduction of Target Mission Concept
- Research Status and External Relations
- Brief Outline of the Subsystems
- Recent R&D Activities for Critical Issues
- Future Work and Schedule

